

Uncertainties

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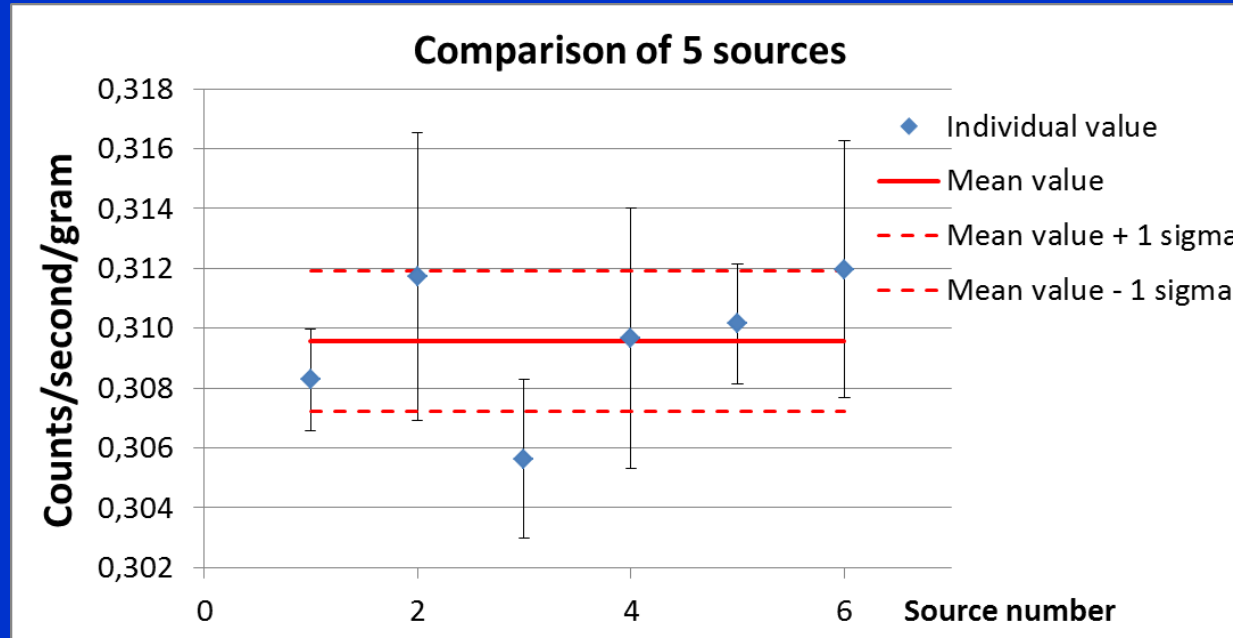
Uncertainties

- The quality of a result obtained by gamma-ray spectrometry (efficiency, activity) depends on:
 - Source (sample) preparation
 - Measurement setup (electronics)
 - Spectrum processing
 - Corrective factors
- Associated uncertainties should reflect all these aspects

Sample representativity

Example: point source from a standard solution to determine efficiency

5 point sources (^{109}Cd) with masses m_1 , m_2 , m_3 , m_4 , m_5



Measuring each source -> count rate per second and per gram

Sample representativity

What about environment samples ? :

What do I wish to measure ?

Is my sample representative of that ?

- Sampling
- Preparation
- Position – Filling height
- Container bottom thickness
 - change the source-to-detector distance
 - attenuation change
- Homogeneity

Electronics

DSP (Digital Signal Processing modules)

Automatic settings

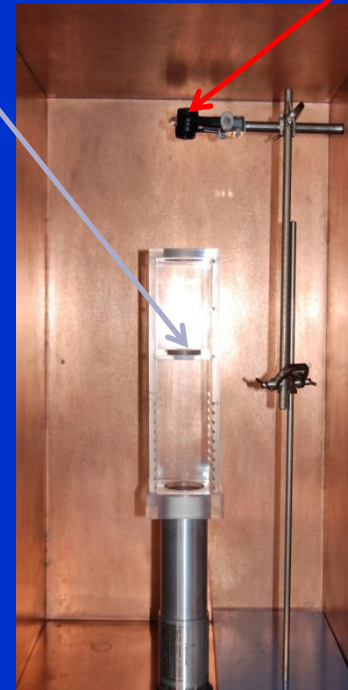
Lot of parameters

^{60}Co +
 ^{137}Cs

^{133}Ba

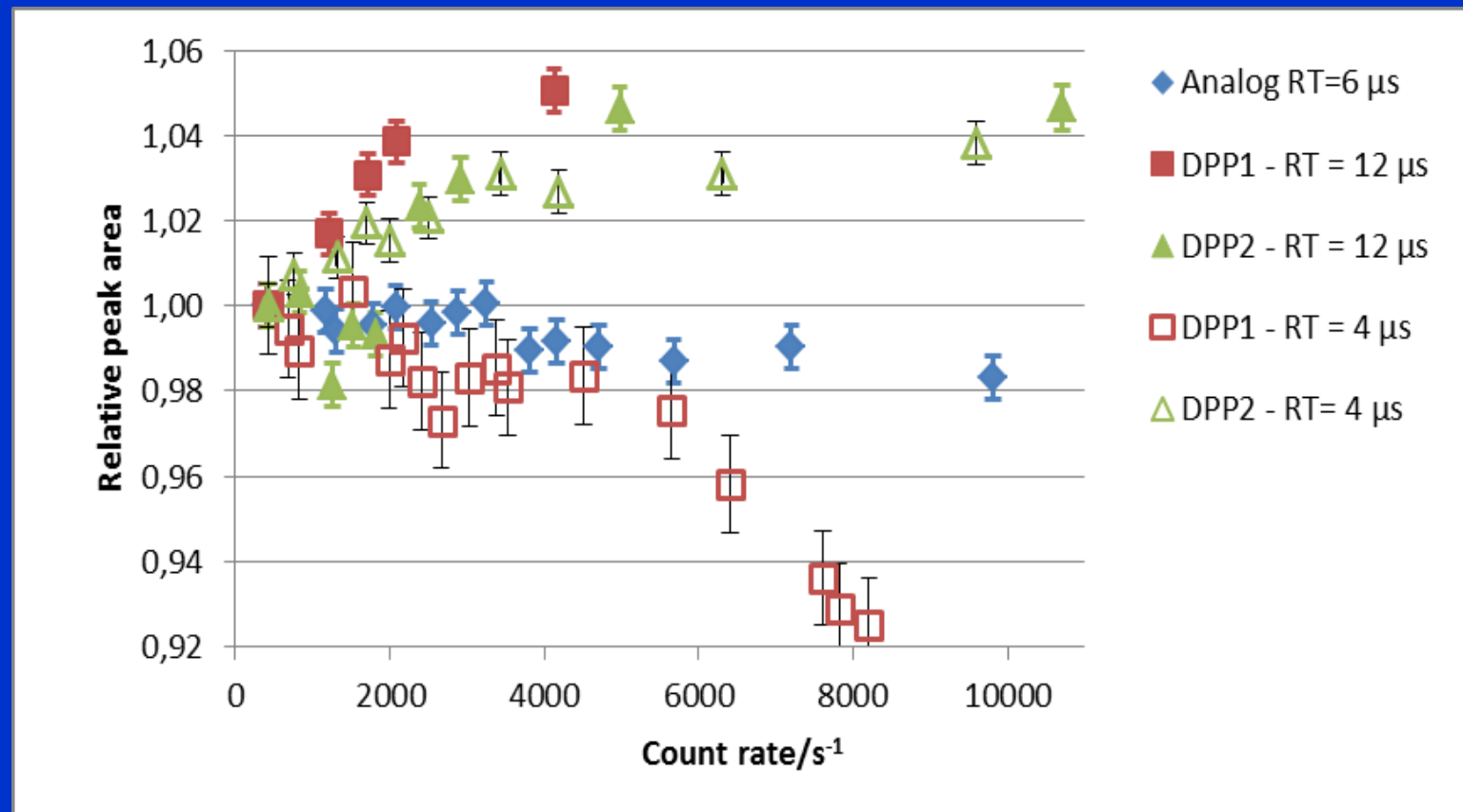
^{60}Co and ^{137}Cs : reference:
fixed position

^{133}Ba : moved close to detector window to
increase the count rate



Tests of electronics

Evolution of the relative 1332 keV peak area versus the count rate for different electronics



Associated uncertainties

- $\varepsilon_i = \frac{N_i \cdot \prod C_{ij}}{A \cdot I_i}$

$$\frac{u(\varepsilon_i)}{\varepsilon_i} = \left[\frac{u^2(N_i)}{N_i^2} + \frac{u^2(A)}{A^2} + \frac{u^2(I_i)}{I_i^2} + \sum_j \frac{u^2(C_{ij})}{C_{ij}^2} \right]^{1/2}$$

Example: Point source of ^{137}Cs - Efficiency at 661.7 keV

$u(A) / A = 0.5 \%$.

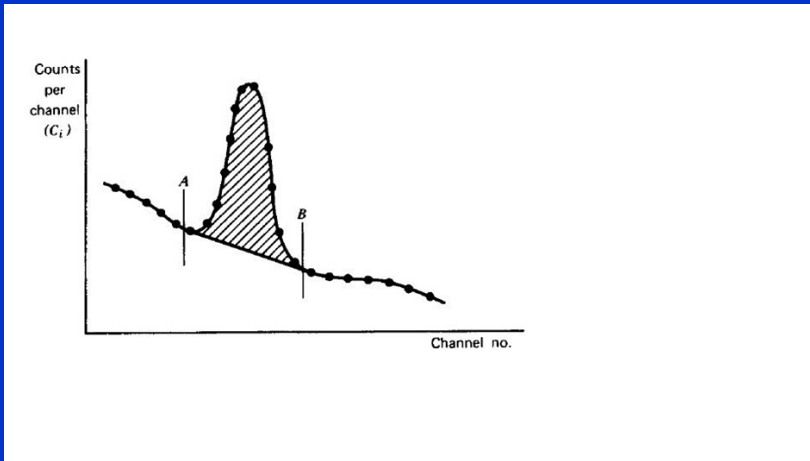
$u(I_i) / I_i : 0.24 \%$.

In the best experimental conditions, where there are no corrective factors, short acquisition time (in comparison with the ^{137}Cs half-life).

Relative uncertainty on peak area (%)	Relative uncertainty on FEP efficiency (%)
1	1.14
0.1	0.56

If $u(A) / A = 0.2 \%$, the FEP efficiency can be obtained with 0.3 % relative uncertainty. This is the minimum that can be experimentally achieved in this very favorable case.

Peak area uncertainty



Depending on
 ‘ the background shape
 the user (definition of binding
channels
 ...

Can be considered as the sum of :

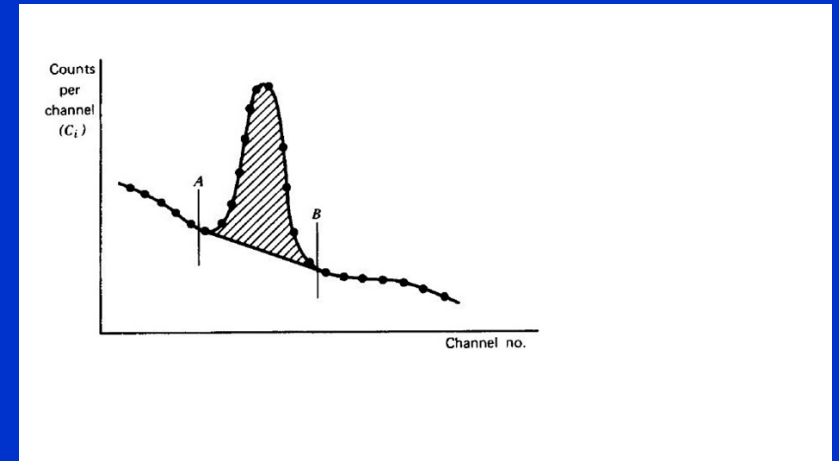
True value of the peak area A_i , and
associated uncertainty (statistics) $A \pm u(A)$
Corrective factor , C_A (=1 if no better data)

$$\frac{u^2(N_i)}{N_i^2} = \frac{u^2(A_i)}{A_i^2} + u^2(C_A)$$

The corrective factor generally cancels out
when the same procedure is used in
calibration and measurement

Peak area determination

- Summing method
- Net = Sum – Background
- $N = N_S - N_B$
- $u(N) = \sqrt{N_S + N_B}$
- Different background shapes

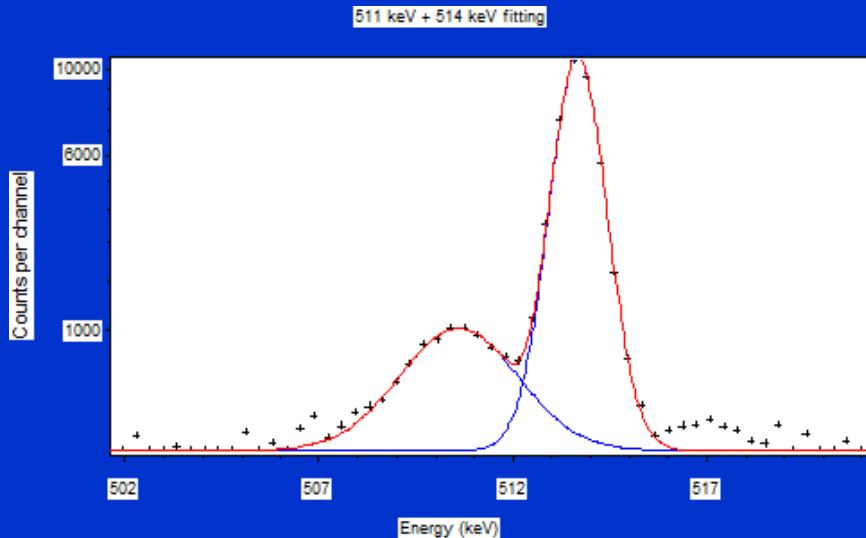


Peak area determination

Fitting method

Using a mathematical method (least squares) to fit a Gaussian to the peak

$$G(E) = G \cdot \exp\left(-\frac{(E - E_0)^2}{2\sigma^2}\right)$$



The total Gaussian area is

$$S(E) = \sqrt{2\pi} \sigma \cdot G$$

with the associated uncertainty:

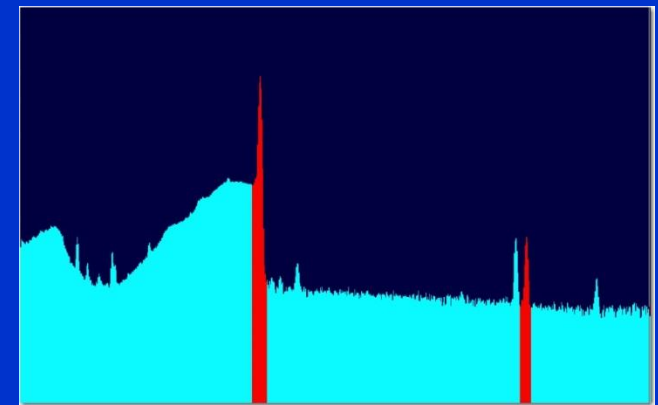
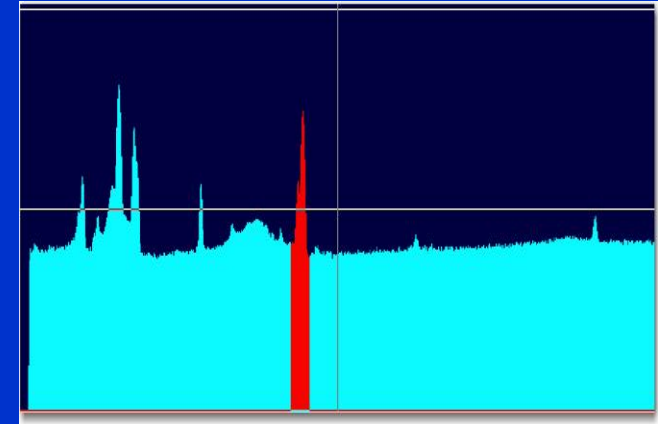
$$\frac{u(S(E))}{S(E)} = \left[\frac{u^2(\sigma)}{\sigma^2} + \frac{u^2(G)}{G^2} \right]^{1/2}$$

$S(E)$ is the Gaussian area integrated over the energy range $[-\infty, +\infty]$. However, 99 % of the area is within the $[E_0 - 2.58 \sigma, E_0 + 2.58 \sigma]$ interval, what can be considered as the practical peak area.

Peak area determination

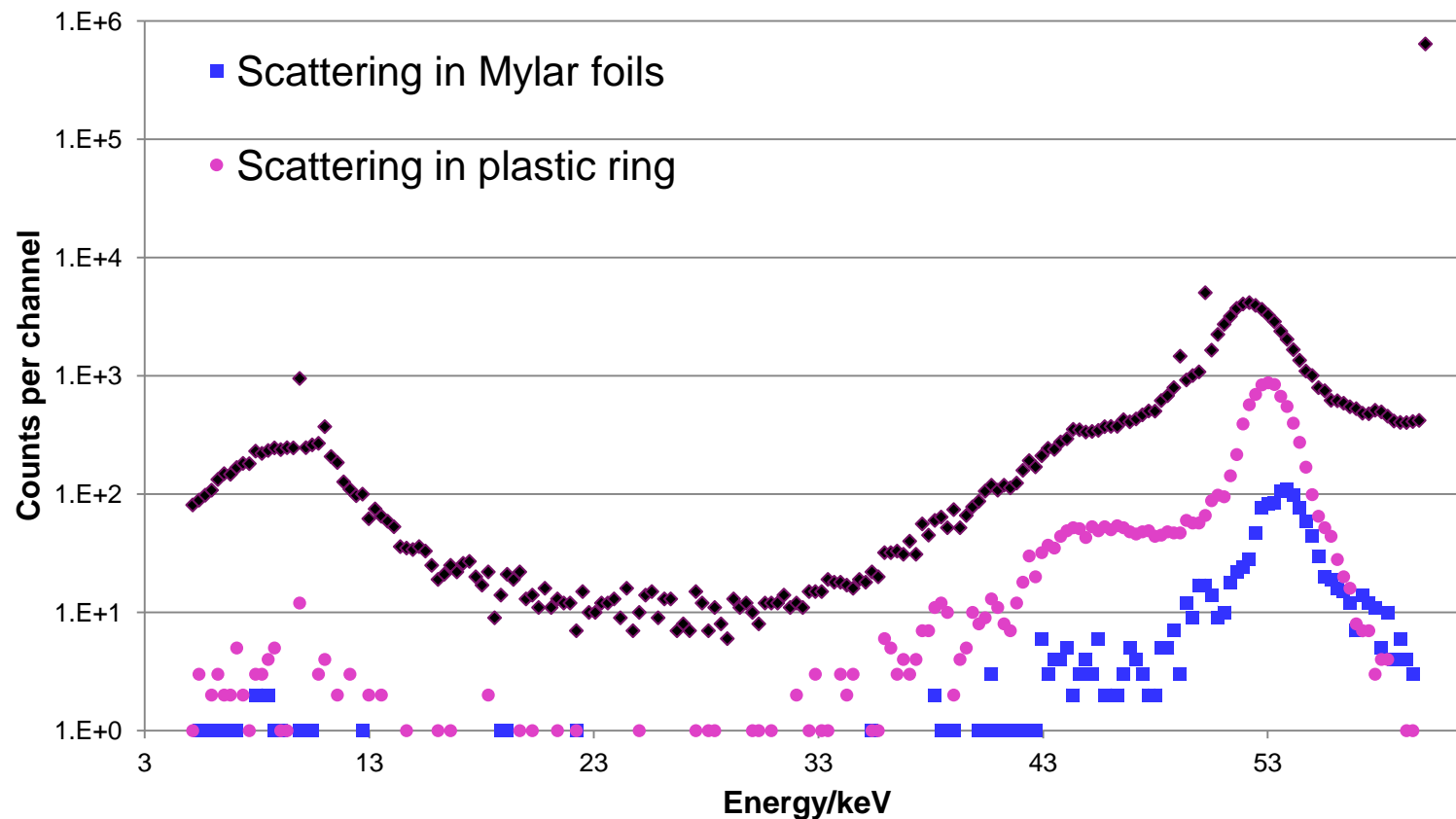
- Scattering effect (low energy range)
- Volume sources

- Example :
 - ^{133}Ba (point source)
 - ^{133}Xe (gaz)
 - Same line at 80 keV

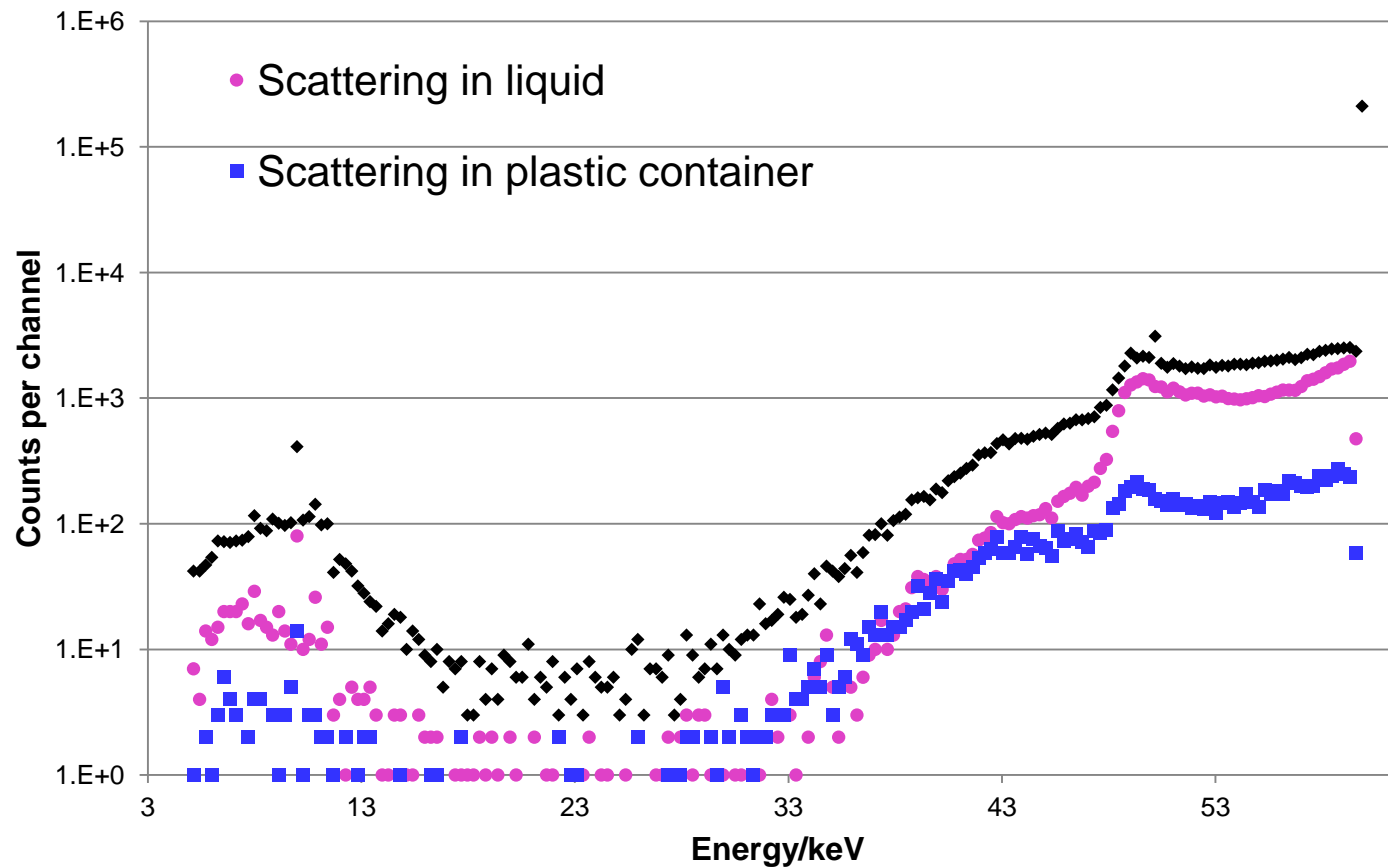


- Efficiency transfer questionable

Monte Carlo simulation for 60 keV photons
Point source at 10 cm



Monte Carlo simulation for 60 keV photons:
Solution (H_2O) in a 50 cm³ plastic container at 10 cm



Corrective factors

- Half-life – decay during measurement
- Attenuation
- Self-attenuation
- Geometry
- Coincidence summing
- Background
- Escape peaks
- Dead-time
- Annihilation in-flight (beta +)

Corrective factors

- Half-life

$$C_T = \exp\left(-\ln(2) \cdot \frac{(T_m - T_r)}{T_{1/2}}\right)$$

- T_m is the measurement time (when the measurement is carried out),
- T_r is the reference time (when the reference activity is known),
- $T_{1/2}$ is the radionuclide half-life.
- It is assumed that the nuclide is not a member of a decay series.

$$\frac{u(C_T)}{C_T} = \ln(2) \cdot \frac{(T_m - T_r)}{T_{1/2}} \cdot \frac{u(T_{1/2})}{T_{1/2}}$$

The uncertainty on the measurement/reference time is generally negligible

Corrective factors

- Decay during measurement (Short half-life)

$$C_{Dec} = \frac{\ln(2) \cdot \frac{t_r}{T_{1/2}}}{1 - \exp\left(-\ln(2) \cdot \frac{t_r}{T_{1/2}}\right)}$$

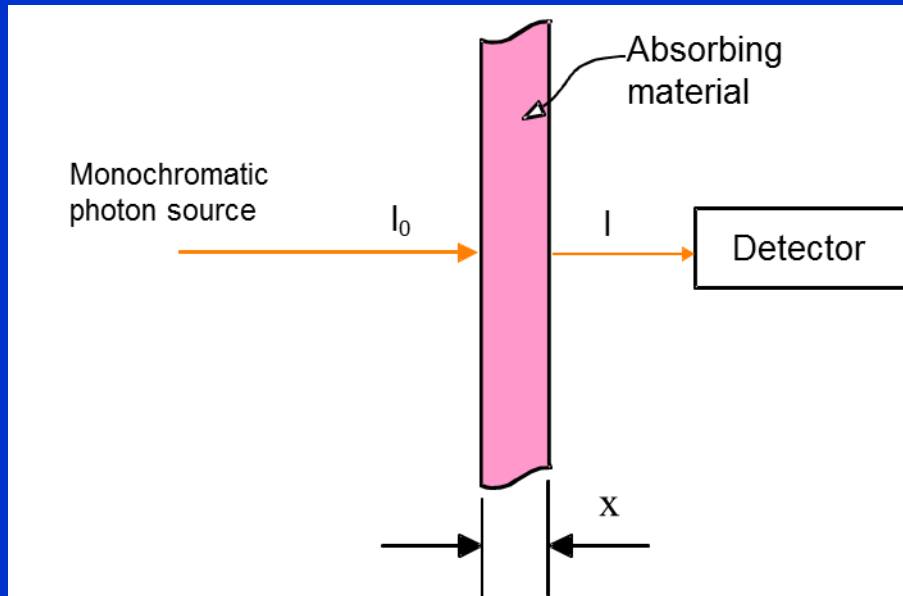
t_r is the real acquisition time.

$$\frac{u(C_{Dec})}{C_{Dec}} = \left[1 + C_{Dec} \cdot \exp\left(-\ln(2) \cdot \frac{t_r}{T_{1/2}}\right)\right] \cdot \frac{u(T_{1/2})}{T_{1/2}}$$

Geometry corrections

- Difference between the calibration and measurement conditions
 - Attenuation (screen)
 - Geometry (distance, filling height)
 - Self-attenuation (matrix, filling height)

Attenuation



$$I(E) = I_0(E) \cdot \exp(-\mu(E) \cdot x)$$

$\mu(E)$ and $\mu/\rho(E)$ are respectively the linear and the mass attenuation coefficients of the screen material and ρ is its density.

$$C_{Att}(E) = \exp(-\mu(E) \cdot x) = \exp\left(-\frac{\mu}{\rho}(E) \cdot \rho \cdot x\right)$$

$$\frac{u(C_{Att}(E))}{C_{Att}(E)} = \mu(E) \cdot x \sqrt{\left(\frac{u(x)}{x}\right)^2 + \left(\frac{u(\mu(E))}{\mu(E)}\right)^2}$$

Valid only for monochromatic photons arriving under normal incidence on the absorbing layer.

Self-attenuation

- Integrating the Beer-Lambert law:
- Simplified correction

$$C_{Self}(E) = \frac{1 - \exp(-\mu(E) \cdot x)}{\mu(E) \cdot x}$$

$$\frac{u(C_{Self}(E))}{C_{Self}(E)} = \left| \frac{\exp(-\mu(E) \cdot x)}{C_{Self}(E)} - 1 \right| \cdot \sqrt{\left(\frac{u(x)}{x} \right)^2 + \left(\frac{u(\mu(E))}{\mu(E)} \right)^2}$$

Approximation valid for small volume and large source-to-detector distance

Practical tools for geometry corrections

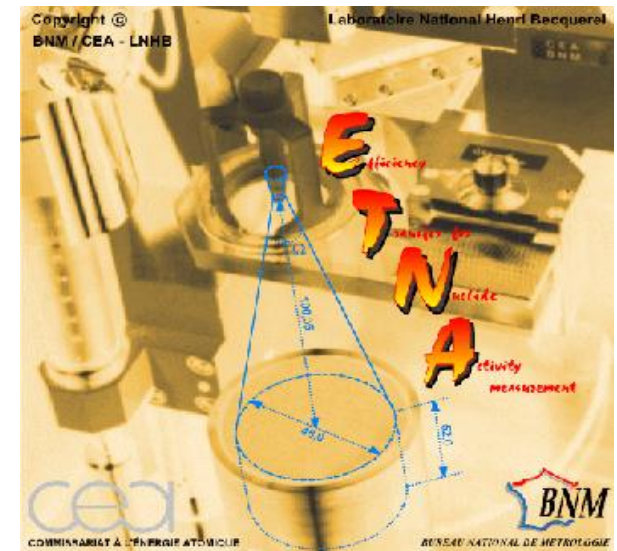
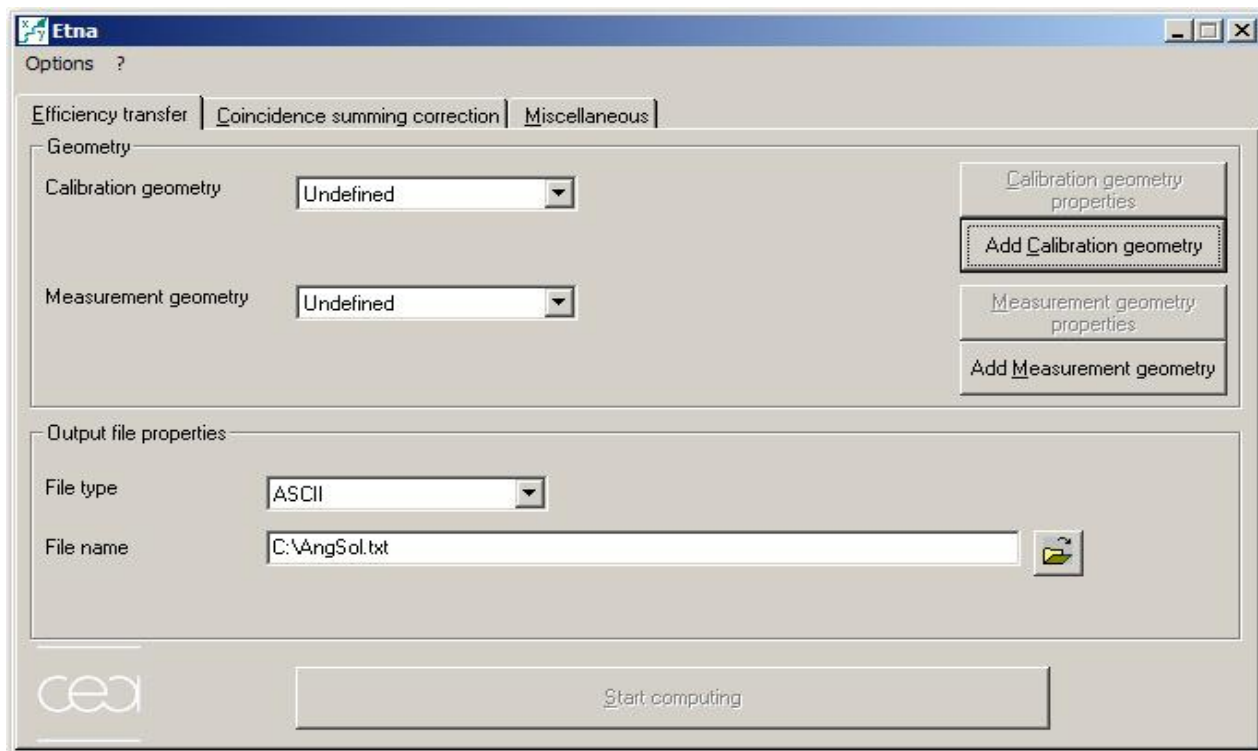
- Change of geometry -> change of efficiency
- Efficiency transfer corrections
- Pure Monte Carlo methods
- GESPECOR, LabSOCS™ (commercial)
- Numerical methods (Moen's principle)
dedicated software such as EFFTRAN,
ANGLE (commercial) or ETNA

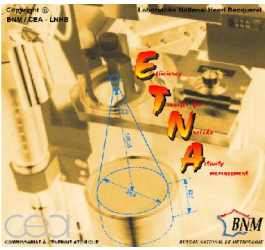
ETNA

(Efficiency Transfer for Nuclide Activity measurement)

ETNA is a software for computing efficiency transfer and coincidence summing corrections for gamma-ray spectrometry.

The software has been developed at the Laboratoire National Henri Becquerel and is available upon request.

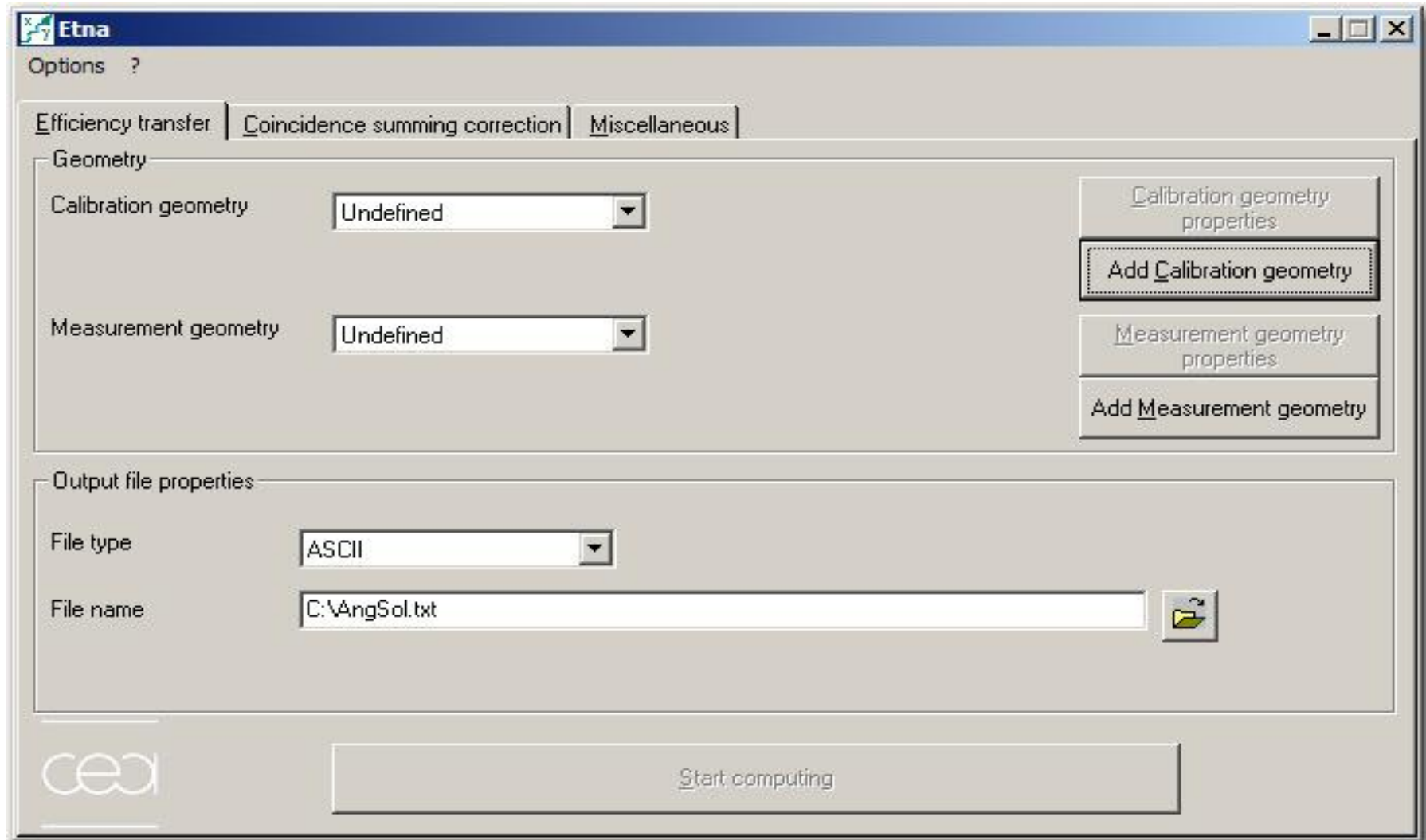




ETNA

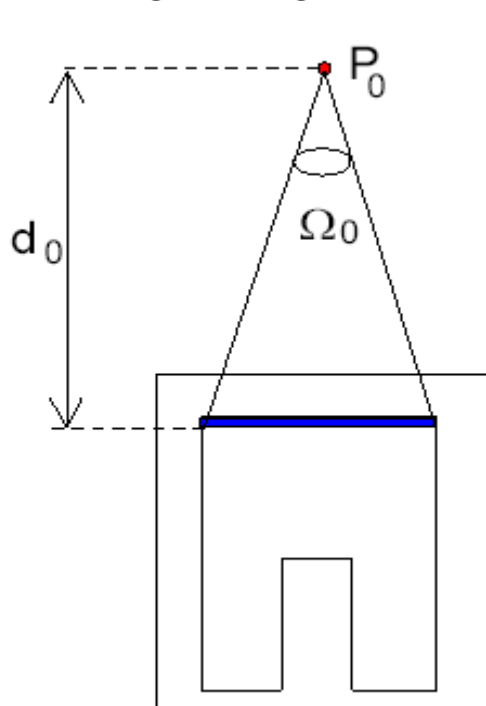
- Transfer of efficiency
 - Semi-empirical method (from a reference efficiency)
 - Coaxial cylindrical geometry (point. disk. cylinder. Marinelli)
- Coincidence summing corrections
 - Knowledge of the efficiency (total and full-energy peak)
 - Possibility of efficiency transfer
 - Decay scheme from Nucleide
- Data management
 - Decay scheme
 - Attenuation coefficients

ETNA main window

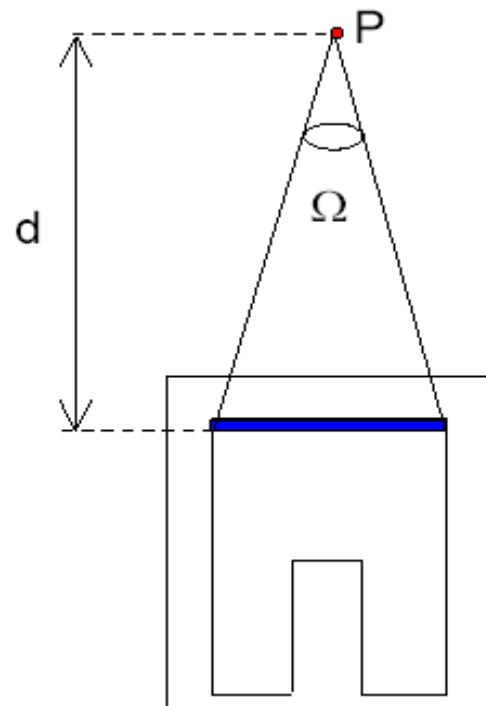


Efficiency transfer principle

Point source moving along the detector axis



$$\mathcal{E}(E, P_0) = \mathcal{E}I(E) \cdot \Omega(P_0)$$



$$\mathcal{E}(E, P) = \mathcal{E}I(E) \cdot \Omega(P)$$

$$\mathcal{E}(E, P) = \mathcal{E}(E, P_0) \cdot \frac{\Omega(P)}{\Omega(P_0)}$$

Solid angle for point source

Using polar coordinates. the solid angle $\Omega(P)$ between point P (r, ϕ, z_s) and the detector entrance surface (disc) is:

$$\Omega(P) = 2 \cdot z_s \int_0^\pi d\phi \int_0^{R_D} \frac{R \cdot dR}{\left[R^2 - 2 \cdot R \cdot r \cdot \cos \phi + r^2 + z_s^2 \right]^{3/2}}$$

R_D is the detector radius.

The geometrical factor should include:

- attenuation in different absorbing layers (air, window, dead layer. ...) : F_{att}

$$F_{att} = \exp \left(- \sum_{i=1}^m \mu_i \cdot \delta_i \right)$$

- absorption in the detector active volume : F_{abs}

$$F_{abs} = f_1 + f_2 \cdot f'$$

$$f_1 = 1 - \exp(-\mu_D \cdot \delta_{1D}) \quad f_2 = 1 - \exp(-\mu_D \cdot \delta_{2D}) \quad f' = \exp(-\mu_D \cdot (\Delta + \delta_{1D}))$$

Solid angle for a cylindrical source

- For a volume source (cylindrical symmetry : radius R_S , thickness H_S , vertical position Z_S):

$$\Omega = \frac{4}{R_S^2 \cdot H_S} \int_{Z_S}^{Z_S+H_S} h \cdot dh \int_0^{R_S} r \cdot dr \int_0^\pi d\varphi \int_0^{R_D} \frac{R \cdot dR}{\left[R^2 - 2 \cdot R \cdot r \cdot \cos \varphi + r^2 + h^2 \right]^{3/2}}$$

- Fatt and F abs must be included in the integration procedure

Integration are numerically performed using the Gauss-Legendre method.

Point sources, discs, cylinders and Marinelli (along the detector axis) are considered.

Input of data

- Requires
 - Detector parameters
 - Source parameters
 - Container
 - Matrix
 - Geometry conditions (source-to-detector distance, screen)
 - Reference efficiency
- Recorded in the « user » database

Efficiency transfer window

Etna

Options ?

Efficiency transfer | Coincidence summing correction | Miscellaneous

Geometry

Calibration geometry: G9 SP a 10 cm *G91*

Measurement geometry: G9 SG15 a 10 cm *G92*

Calibration geometry properties

Add Calibration geometry

Measurement geometry properties

Add Measurement geometry

Output file properties

File type: ASCII

File name: C:\AngSol.txt

cea

Start computing

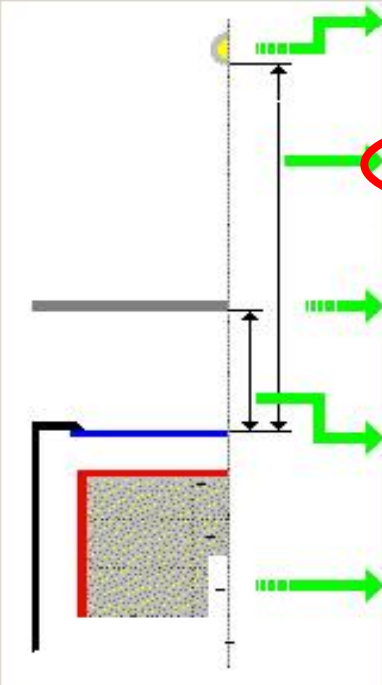
Input of geometry parameters

Calibration geometry properties

Geometry
Geometry Comment
☒ Calibration geometry

Creation date Last modification date

Geometry properties



Source
Reference

Source-detector distance ± mm

Measurement environment

Absorber

Absorber-detector distance ± mm

Detector

Efficiency transfer results

```
ETNA _____ Version 5.5 Rev 55

Filename : C:\Documents and Settings\LEPY.BECQUEREL\Bureau\test

jeudi 23 septembre 2010
Processing identification : Efficiency transfer
  Calibration geometry : G9 SP at 10 cm (G91)
  Calibration source : Point source (Reference)
  Calibration source - detector distance : 103.4 mm
  Calibration absorber : None
  Calibration absorber - detector distance : 0 mm
  Measurement geometry : G9 SG15 a 10 cm (G92)
  Measurement source : Volume SG15 (Filled with HCl 1N 41 mm)
  Measurement source - detector distance : 104 mm
  Measurement absorber : None
  Measurement absorber - detector distance : 0 mm
  Detector : G9

Results :
Energy      Calibration      Measurement      Ratio
(keV)       efficiency        efficiency
00020.000   00.00288000       00.00058685     00.20376736
00050.000   00.00812000       00.00391170     00.48173645
00080.000   00.00932000       00.00482243     00.51742811
00100.000   00.00924000       00.00489268     00.52951082
00120.000   00.00882000       00.00476234     00.53994785
00150.000   00.00796000       00.00438473     00.55084548
00200.000   00.00655000       00.00371224     00.56675420
00250.000   00.00543000       00.00313296     00.57697238
00300.000   00.00458000       00.00268981     00.58729476
00400.000   00.00346000       00.00208513     00.60263873
00500.000   00.00278000       00.00170495     00.61329137
00750.000   00.00191000       00.00120835     00.63264398
01000.000   00.00149000       00.00096245     00.64593960
01250.000   00.00123000       00.00080666     00.65582114
01500.000   00.00105000       00.00069649     00.66332381
01750.000   00.00089500       00.00059884     00.66909497
02000.000   00.00076600       00.00051653     00.67432115

CEA / LNE-LNHB _____
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Coincidence summing

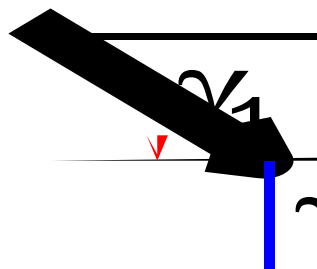
- Effect due to the decay scheme
- Even at low counting rate
- More important at short source-to-detector distance
- Same kind of tools as for Efficiency transfer
 - Monte Carlo
 - GESPECOR
 - Dedicated numerical (ETNA, other ?)

Calculation principle

- ETNA uses a numerical method, according to Andreev, Mc Callum principle:

zX

β



P_{12} : probability for emitting γ_2 simultaneously with γ_1
 ε_i : FEP efficiency for energy E_i
 η_i : Total efficiency for energy E_i

$$C_1 = \frac{1}{1 - P_{12} \cdot \eta_2}$$

$$C_2 = \frac{1}{1 - P_{21} \cdot \eta_1}$$

$$C_3 = \frac{1}{\left(1 + \frac{I_{\gamma 1}}{I_{\gamma 3}} \cdot \frac{\varepsilon_1 \cdot \varepsilon_2}{\varepsilon_3} \cdot P_{12} \right)}$$

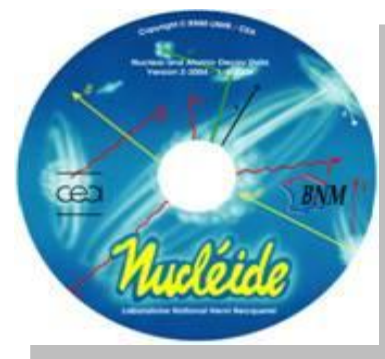
Calculation principle (2)

- Double coincidences
- Coincidences with K X-rays (electron capture or internal conversion) are computed
- Correction for K-X-rays (from gamma or X rays) are computed
- Beta+ emitting nuclides are considered (modification of the decay scheme)
- No angular correlation

ETNA – Input data

ETNA requires:

1. Decay scheme (Nucleide database)
2. FEP and total efficiency for at least one source-to-detector geometry («calibration geometry » recorded in the « user » database)



ETNA – Coincidence tab

Options ?

Efficiency transfer | Coincidence summing correction | Miscellaneous

Nuclide: Ba133 Daughter nuclide: Cs133

Geometry

Calibration geometry: G1 SP reference Source ponctuelle à 10 cm

Measurement geometry: G1 SP reference Source ponctuelle à 10 cm

☐ Measurement geometry different from calibration geometry

Calibration geometry properties

Add Calibration geometry

Measurement geometry properties

Add Measurement geometry

Output file properties

File type: ASCII

File name: C:\Corco.txt

☒ Simplified computing ☐ Complete computing

Start computing

CEA BNM

Coincidence correction results

- dimanche 22 février 2009
- ETNA _____Version 5.5 Rev 51
- Filename :C:\Documents and Settings\ML118236\Bureau\Workshop_ICRM\Presentations\ETNA\test_ETNA

- dimanche 22 février 2009
- Processing identification : Coincidence summing correction (simplified computing)
- Nuclide :Ba133
- Daughter nuclide :Cs133
- Half-life threshold :0.000001 s
- Calibration geometry : G1 SP reference (Source ponctuelle à 10 cm)
- Calibration source :Source ponctuelle
- Calibration source - detector distance :100 mm
- Calibration absorber :None
- Calibration absorber - detector distance :0 mm
- Measurement geometry :Calibration geometry
- Detector :G1 - pièce 6A

- Results :

- Error codes : 0 0
- X-ray correction : 01.015880

Starting level	Arrival level	Energy (keV)	Gamma-gamma correction	Gamma-X correction	Total
004	003	00053.162	01.013962	01.010219	01.024324
002	001	00079.614	01.015207	01.012325	01.027720
001	000	00080.998	01.011478	01.007984	01.019554
002	000	00160.612	00.993490	01.007235	01.000678
003	002	00223.237	01.009461	01.019791	01.029439
004	002	00276.399	01.008560	01.015827	01.024522
003	001	00302.851	01.005028	01.015414	01.020519
004	001	00356.013	01.003565	01.011468	01.015074
003	000	00383.849	00.991597	01.010308	01.001818

- :

- CEALNHB _____BNM

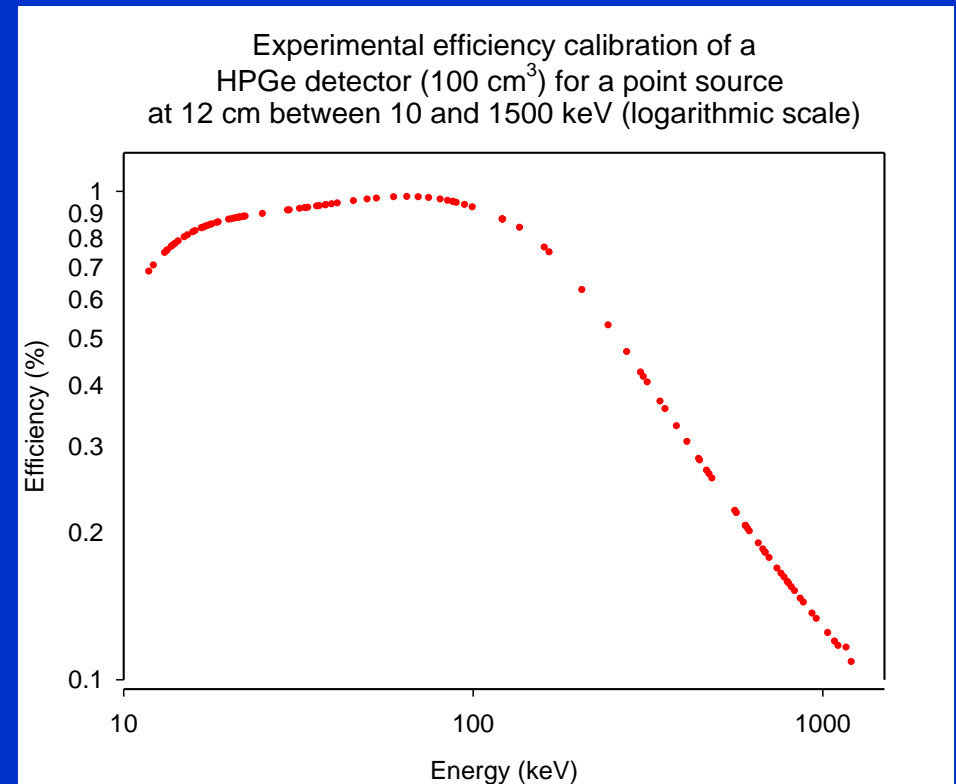
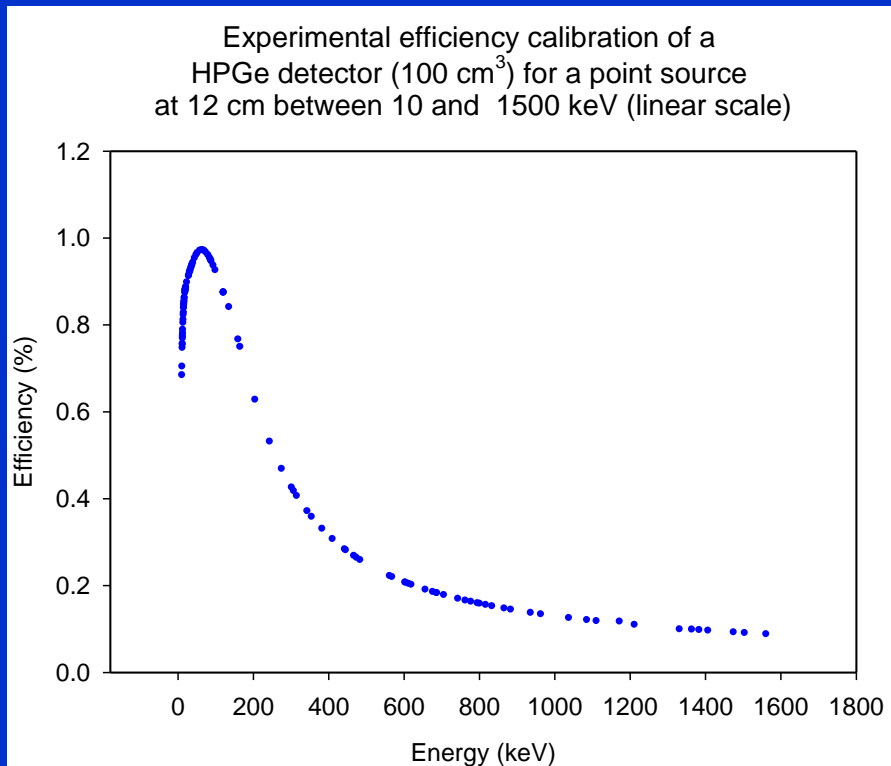
Other possible corrections

- Escape peaks
 - Annihilation (-511 and – 1022 keV)
 - Ge K X-rays (-11keV)
- Background
 - Natural radionuclides
 - Variation of Rn content versus time
- Dead time ?(if the correction is not accurate depending on the counting rate)

EFFICIENCY CURVE

Efficiency calibration : mathematical fitting (1)

Determination of the best fitted function to a given set of experimental data (energy, efficiency)



In the logarithmic scale , the shape is smoother than in the linear scale.

Efficiency calibration : mathematical fitting

Functions frequently used:

Polynomial fitting in the log-log scale:

$$\ln \varepsilon(E) = \sum_{i=0}^n a_i \cdot (\ln E)^i$$

$$\ln \varepsilon(E) = \sum_{i=0}^n a_i \cdot E^{-i}$$

Remarks :

- a_i coefficients are determined using a least-squares fitting method
- experimental data must be weighted
- the polynomial degree (n) must be adjusted depending on the number of experimental data (p) : $n \ll p$
- in some case two different functions can be used with a cross point
- check the resulting fitted curves !

Efficiency calibration : mathematical fitting

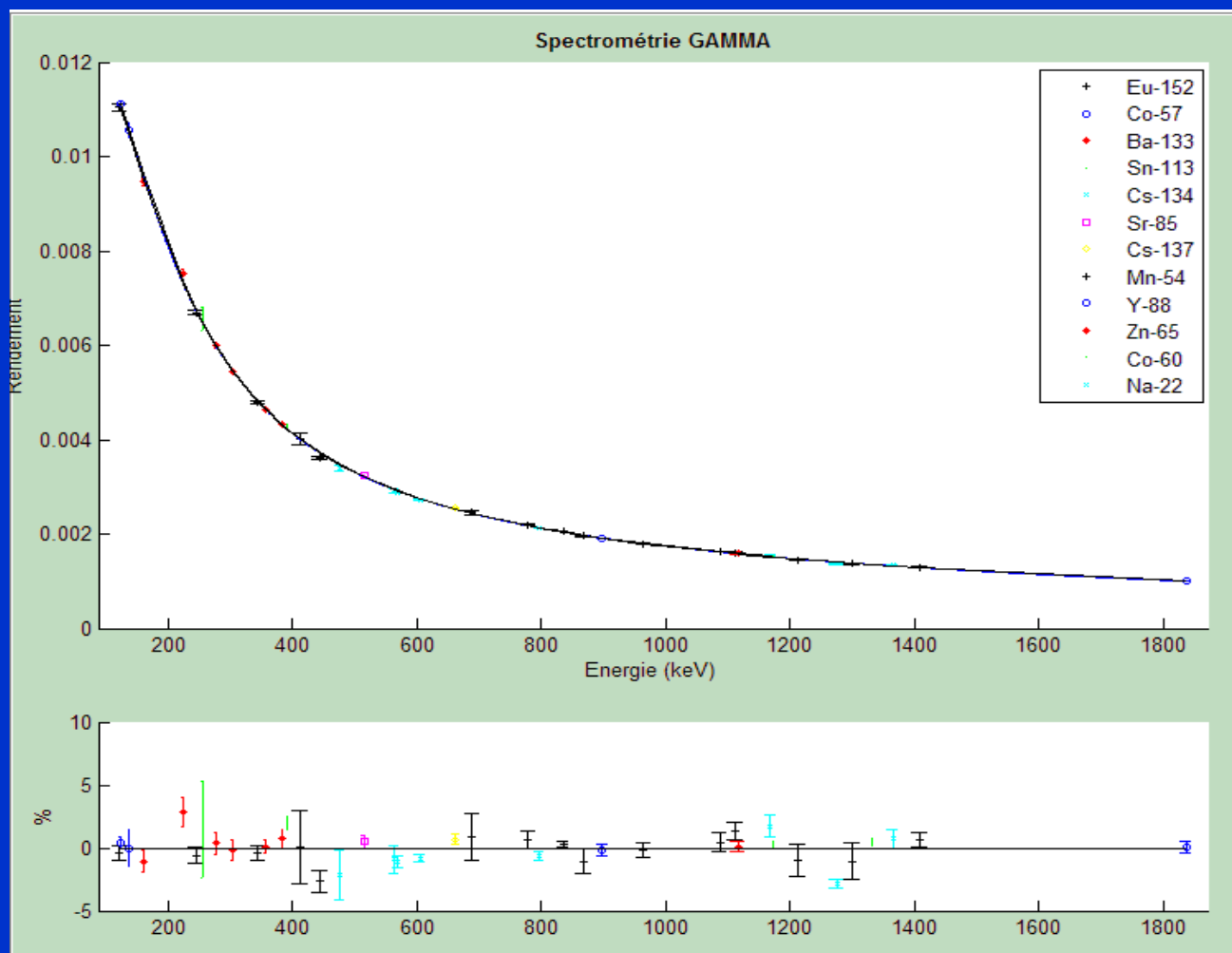
Example : 40 experimental values
in the 122-to-1836 keV range

Fitting function :

$$\ln \varepsilon(E) = \sum_{i=0}^n a_i \cdot (\ln E)^i$$

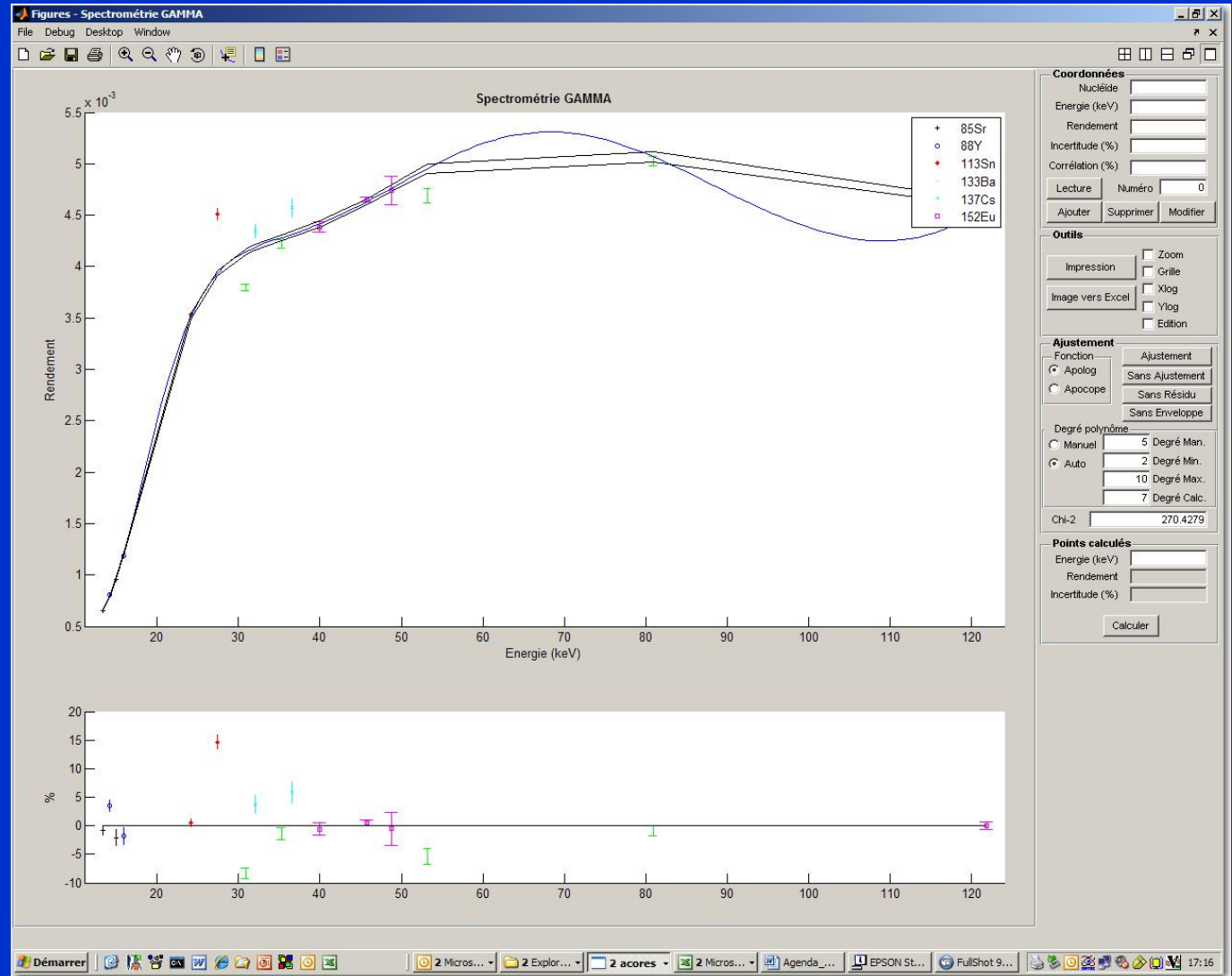
Adjusted coefficients :

fitting	122 to 1836 keV
deg0	-34,11961
deg1	48,16797
deg2	-25,89215
deg3	5,80219
deg4	-0,40503
deg5	-0,01632



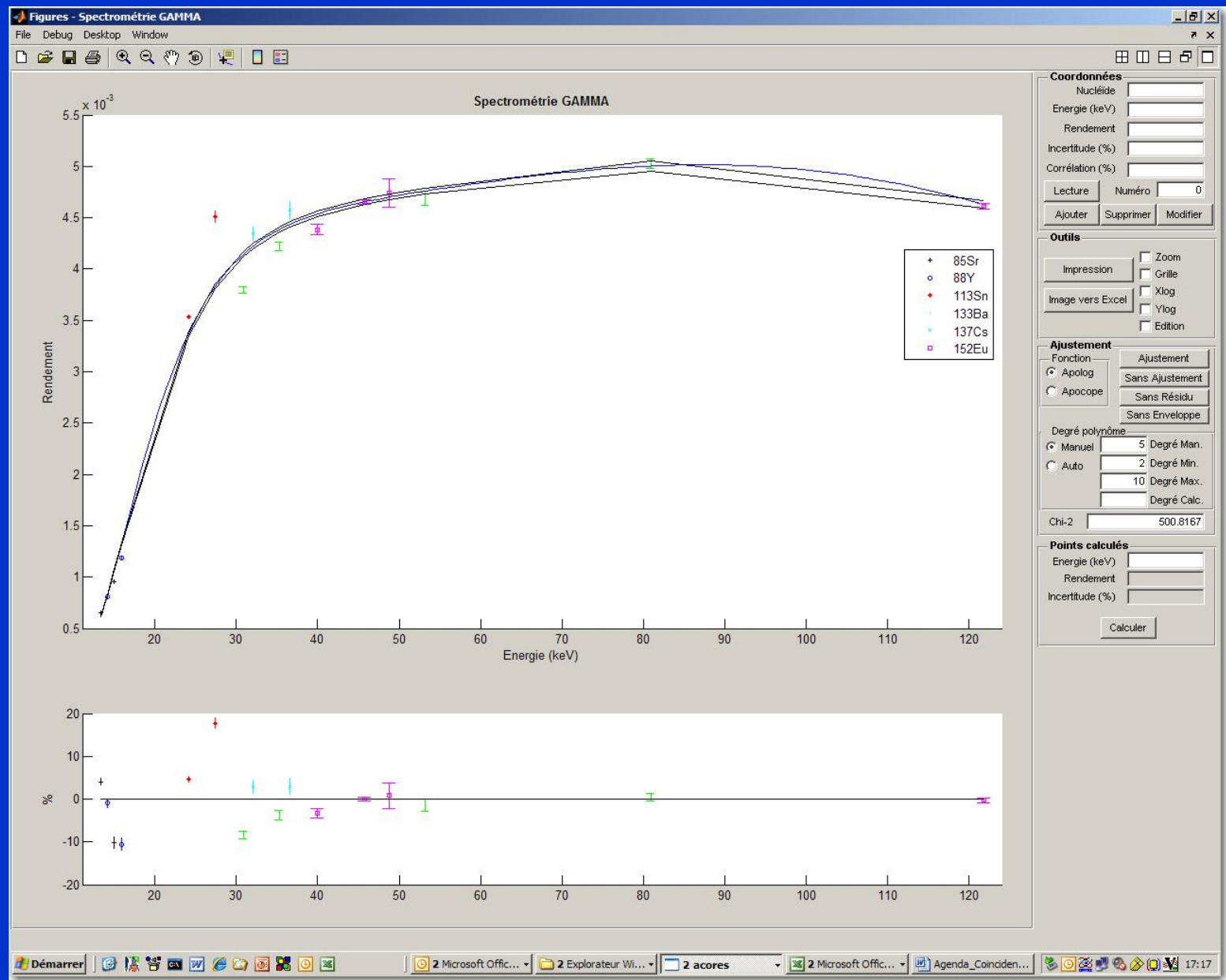
Efficiency calibration : mathematical fitting

Efficiency fitting must be visually checked



In the case
of cross
points
be carefull :

- Avoid
zones with
important
inflexion
- Avoid high
degree
polynomials



Uncertainty on the fitted efficiency

- The individual points have uncertainties
- The mathematical fitting can result in lower uncertainties
- Some correlations exists
 - Input data : one radionuclide- several energies
 - Calibration procedures,
 - etc.
- Careful examination is necessary

Conclusions

- Uncertainties are generally underestimated.
- Important to take each component into consideration.
- Corrective factors should be as close to 1 as possible (experimental conditions).